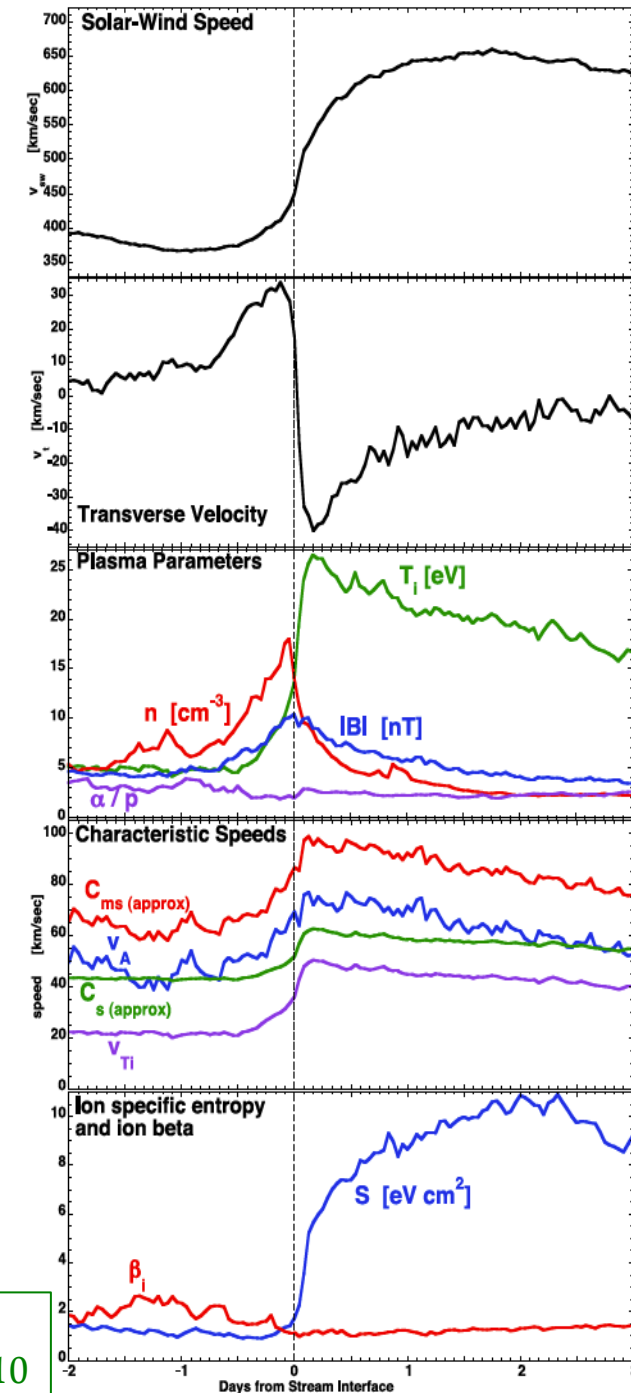
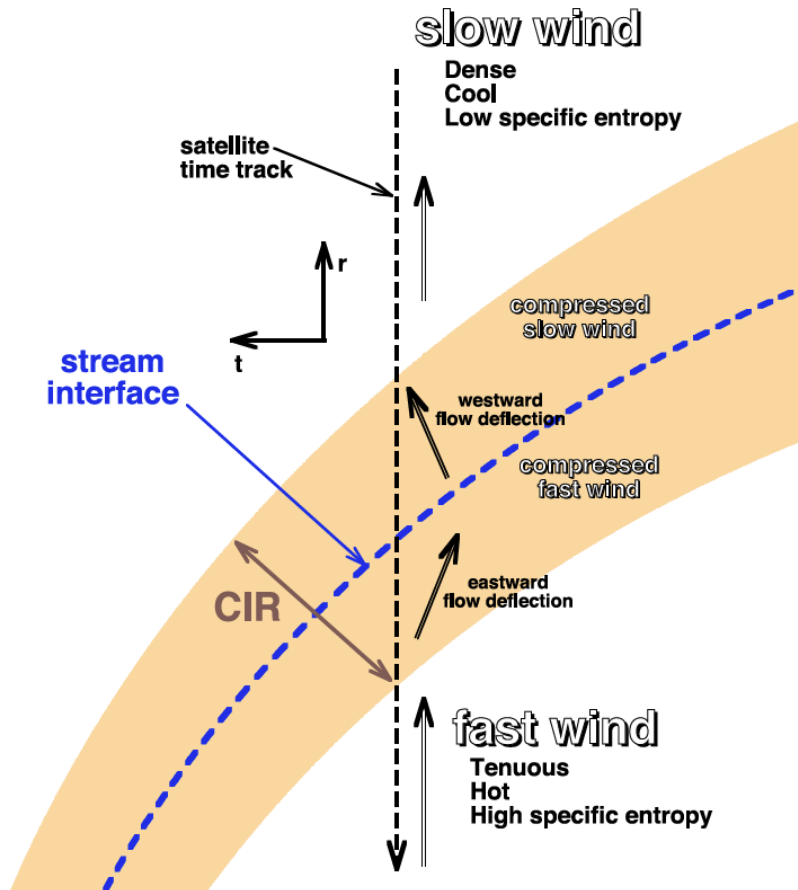


Comparing HELCATS CIR catalogues derived from white-light images and in-situ measurements

Illya Plotnikov¹, Alexis P. Rouillard¹

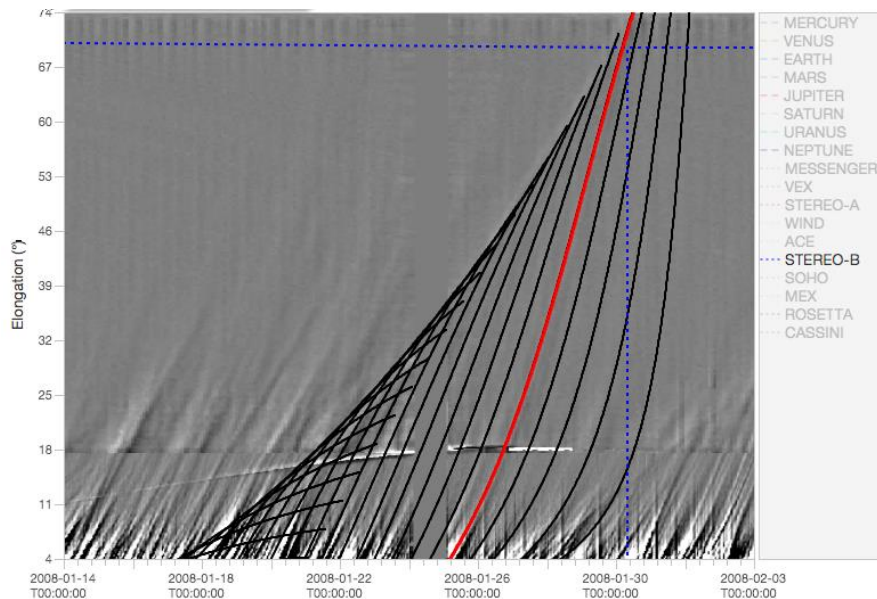
(1) IRAP-CNRS / UPS, Toulouse, France

In-situ view of a CIR



Borovsky & Denton
JGR, 115, A10101, 2010

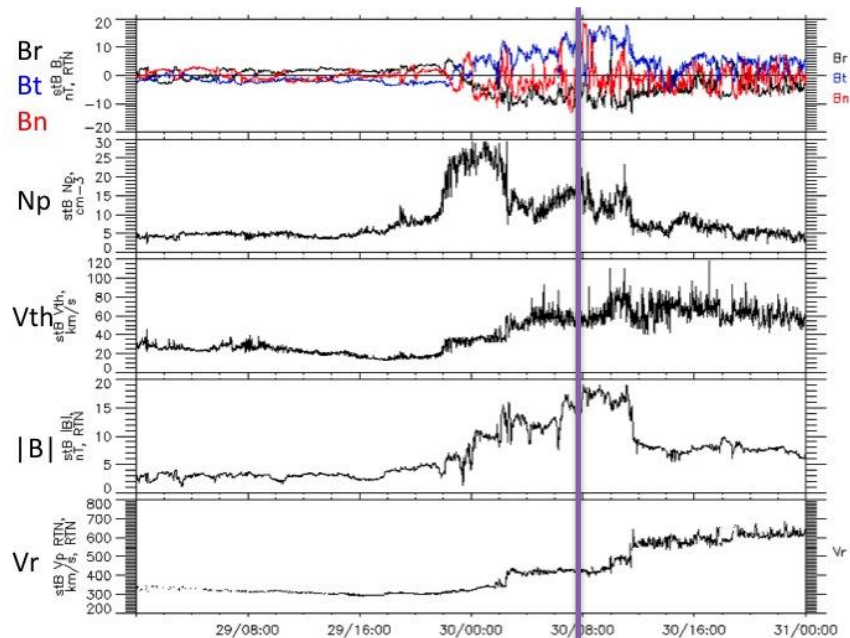
Figure 5. A sketch of a CIR in the vicinity of 1 AU. Standard view in the reference frame of the Sun in RTN coordinates. The reference frame emphasizes the westward-eastward flow deflection (arrows) in the CIR (brown shading).



- Fit of the one density enhancement tblob.
- Reconstruct all traces and the envelope (locus of the enhanced visibility)
- Backproagate to the Sun (CH in proximity?)
- Propagate to different probes (Stereo A-B, ACE, Wind...)

- On 2008-01-24
- Fitted speed: 291 km/s
- No Coronal Hole seen at the footprint of the event.
- Predicted arrival time in the middle of the CIR.

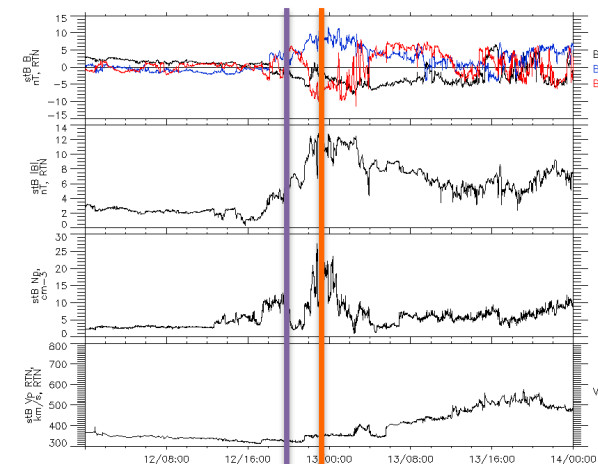
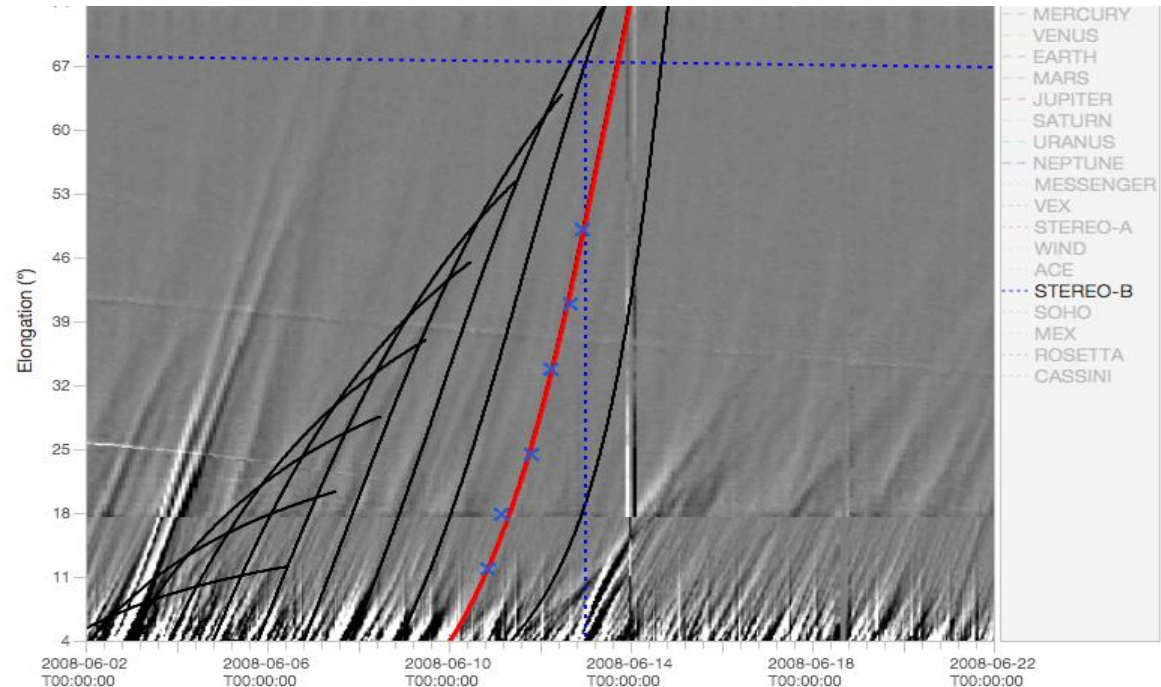
AMDA Stereo B in situ data



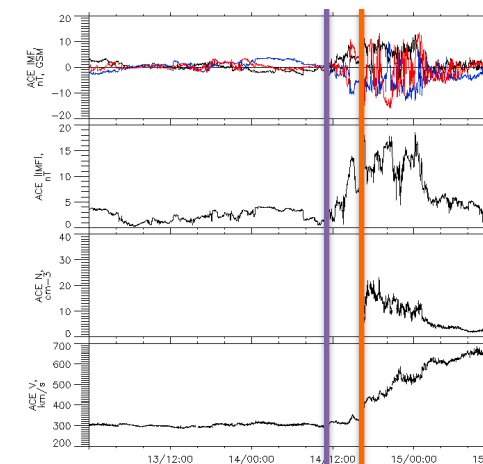
- Question: relevance of the heliospheric imagers to predict SIR (CIRs) arrival time at distant sources?
- A subset of 61 events selected (april 2007 – december 2008)
- Predicted arrival times and speeds at Stereo A-B, Ace, Wind, Rosetta etc. are recorded for every event.
- The incertitude on the velocity is typically about 10%.

Start time: 2008-06-08T21:10:30. Speed: 335 ± 12 km/s

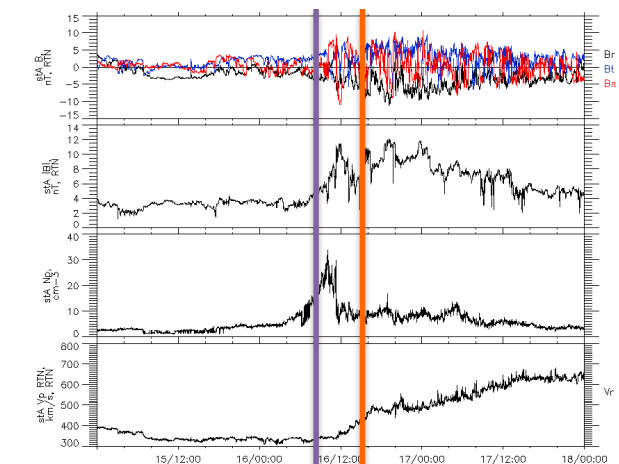
- Stereo B: impact at 2008-06-12T18:58:18
- Ace: impact at 2008-06-14T10:47:00 (Borovsky & Denton 2010: stream interface at 2008-06-14T16:27:00)
- Stereo A: impact at 2008-06-16T10:36:0
- New fit in better agreement with ace data. Close to the stream interface.



Jun 08 Created by AMDA(C) V2.0 Fri May 15 15:23:21 2015

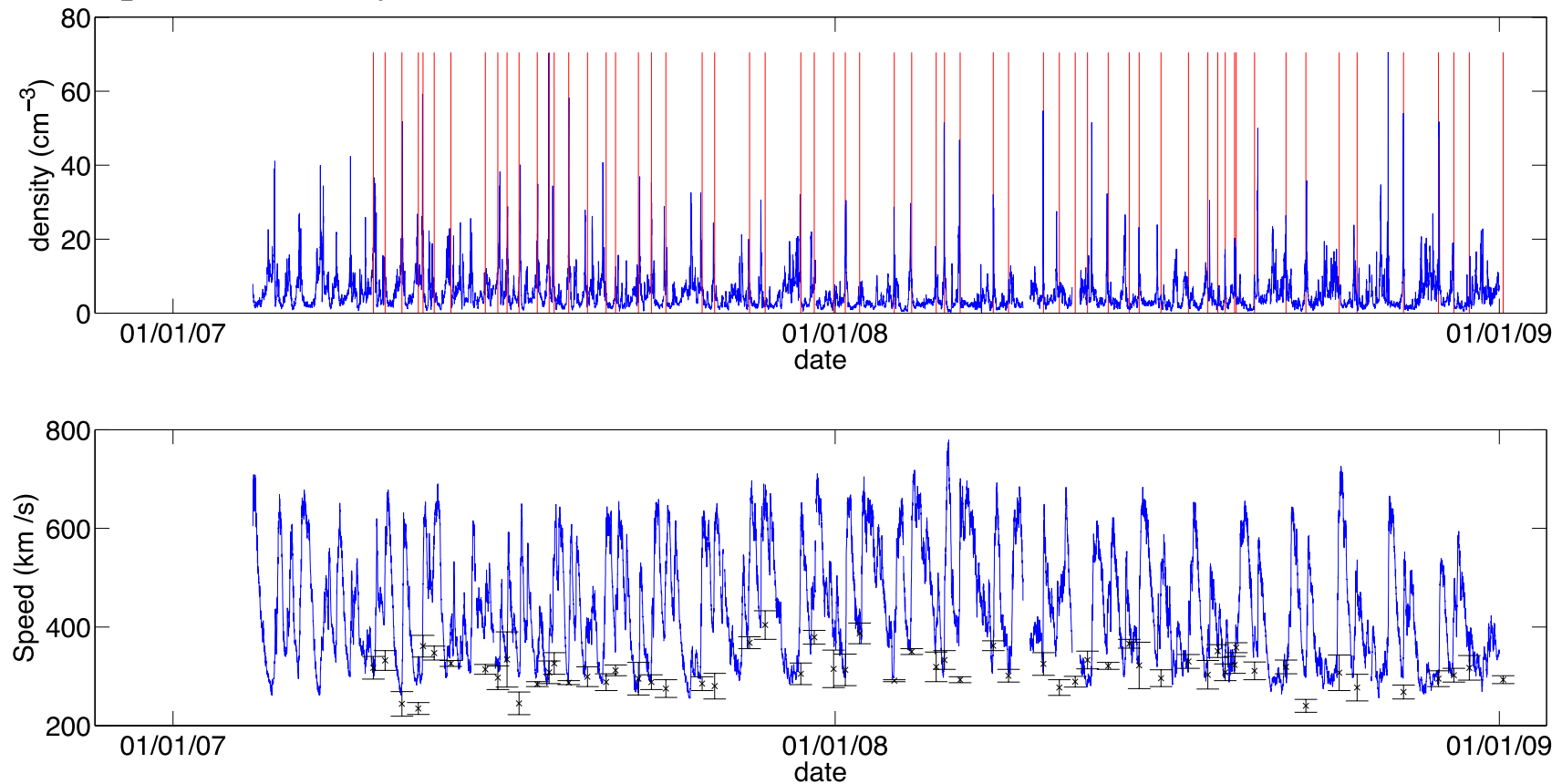


Jun 08 Created by AMDA(C) V2.0



Jun 08 Created by AMDA(C) V2.0 Fri May 15 15:23:57 2015

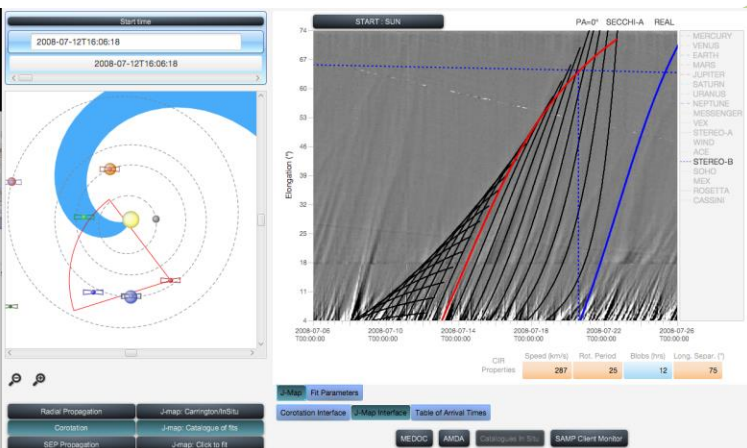
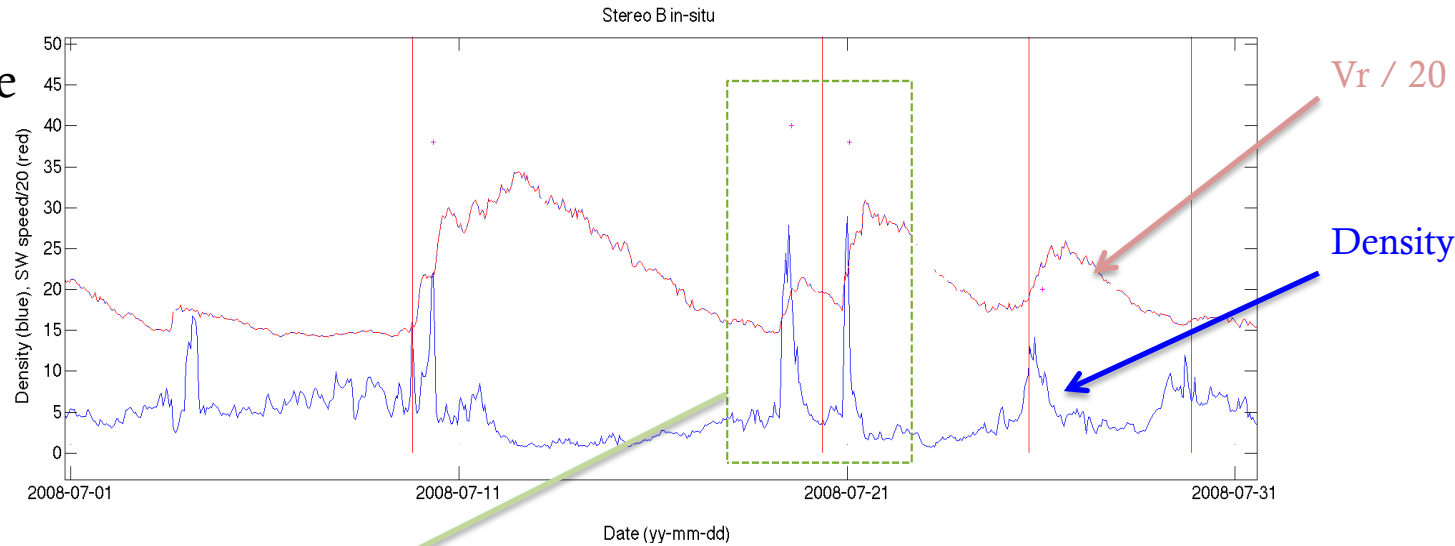
In-situ Stereo A density and radial speed with predicted arrival times (red vertical lines) and speeds (black symbols).



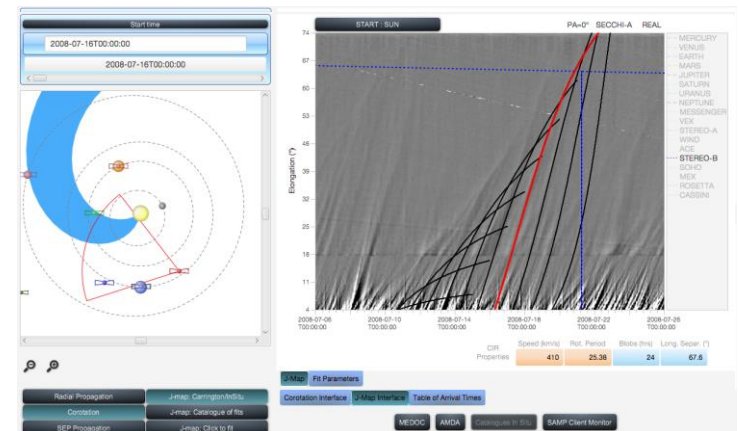
Good correlation with the slow SW but underestimated. Predicted arrivals are generally after the density pics (up to 2 days difference).

Special cases: very close SIRs

Identified as unique
SIR in the first
J-map run

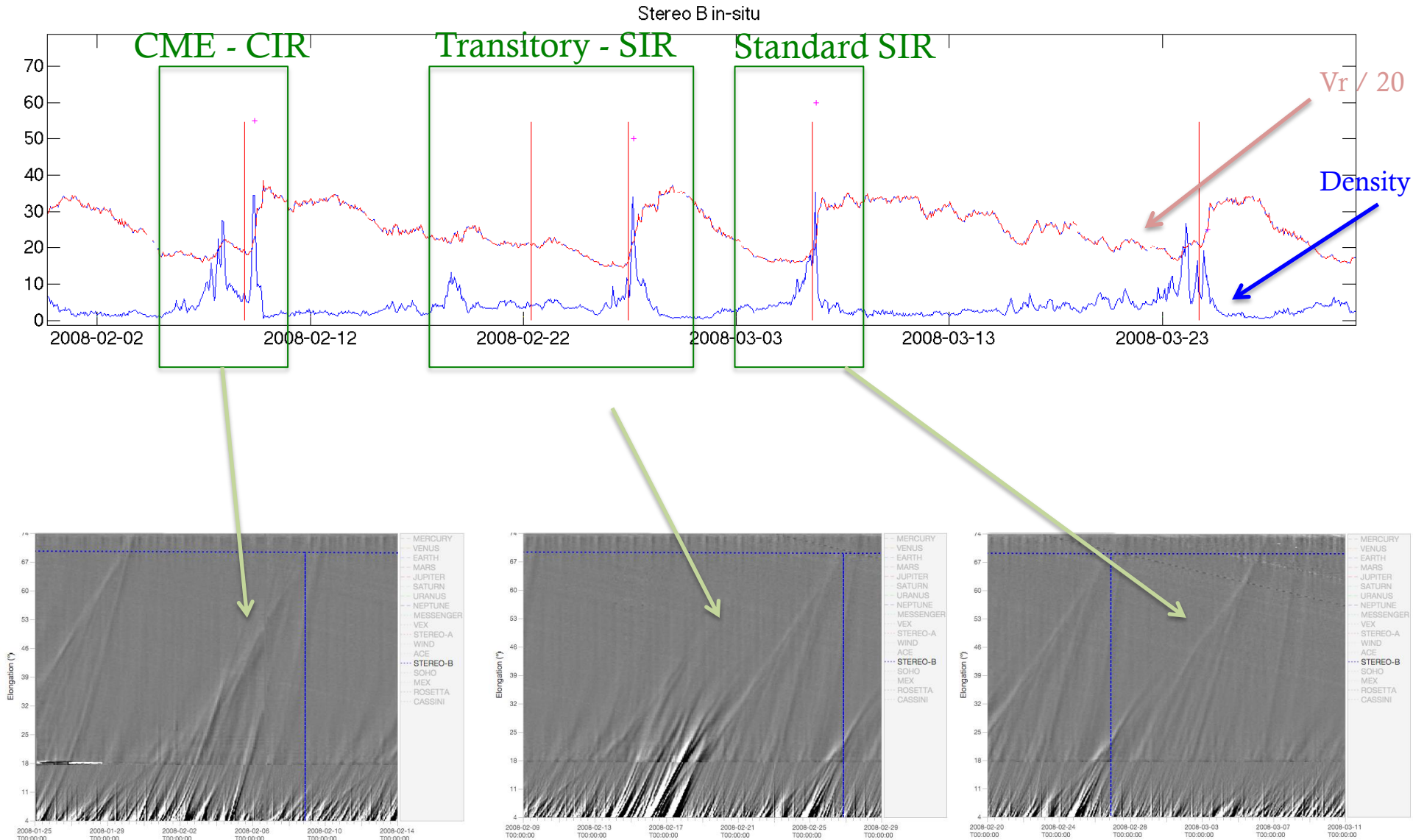


Rectify CH
and speed

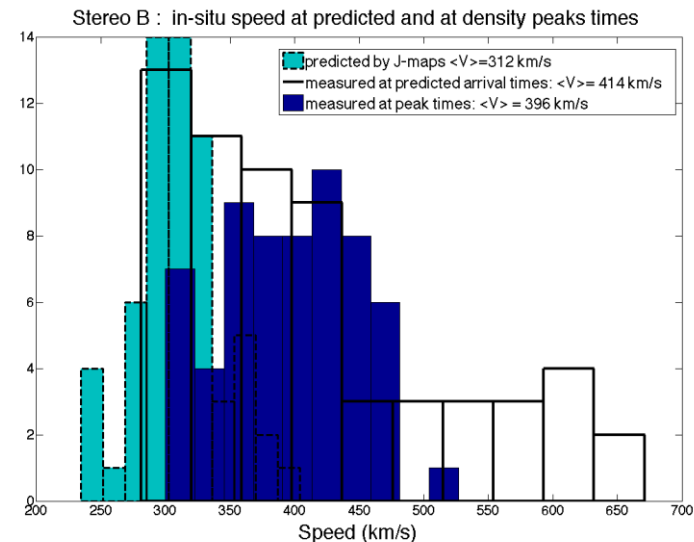
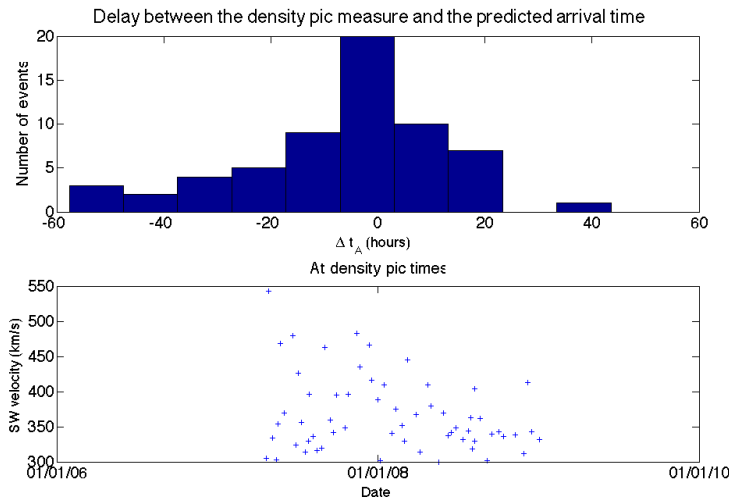


Special cases: CME-SIR, transitory stuff...

February-march 2008 Stereo B data.

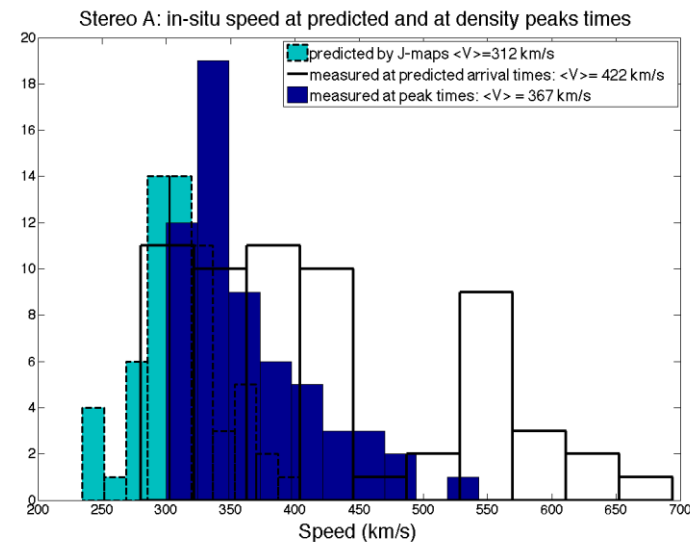
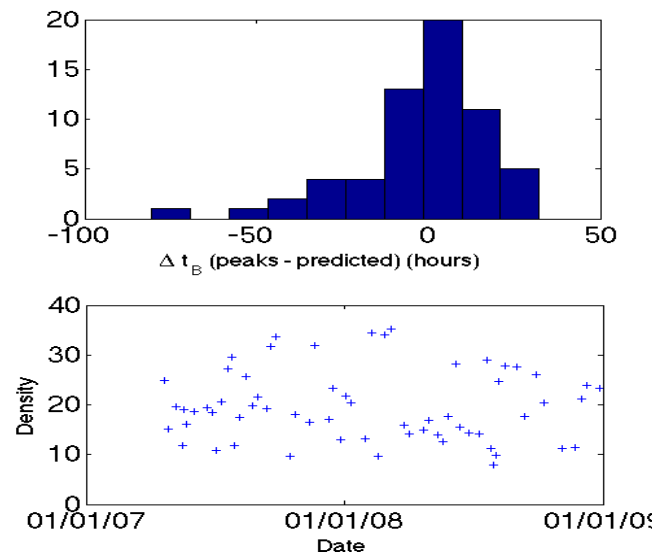


- One by one correlation from predicted arrival times to the closest density pics in the in-situ data for Stereo-A and B.



- Trend to predict after the actual SIR arrival
(max difference = 2 days, mean = 14.4 hours, median=9.3 hours)
- Slow wind speed seems to be underestimated by J-map fits.
Track are followed clearly for $\alpha < 45$ deg.
The speed might be lower at these distances
- Density peaks are not so relevant for slow wind (stream interface).
Need to measure 1-2 days before. (see comparison to L. Jian results).

- One by one correlation from predicted arrival times to the closest density pics in the in-situ data for Stereo-A and B.



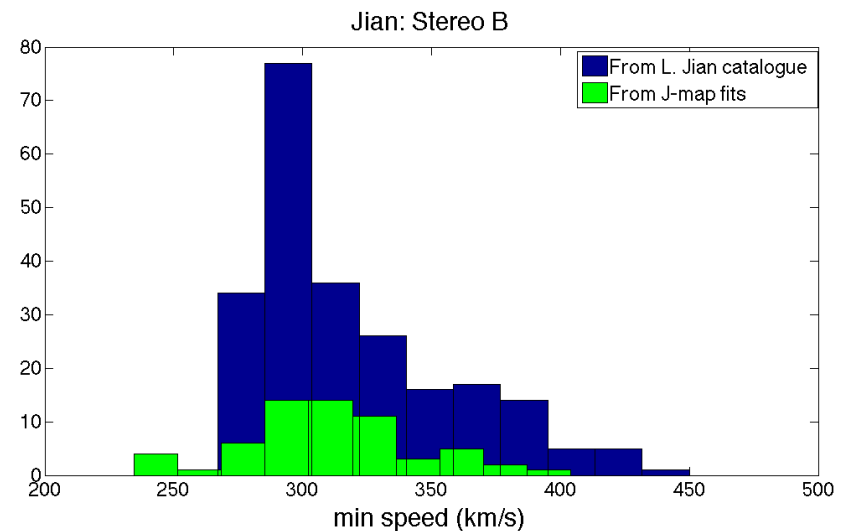
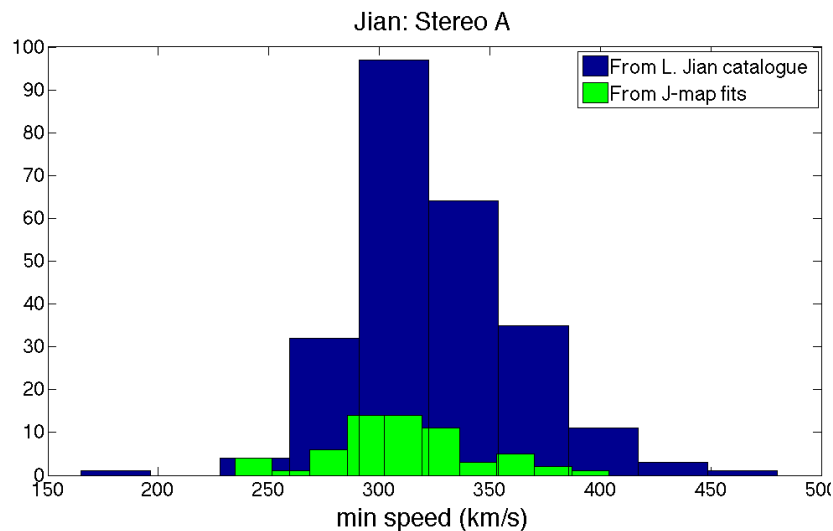
- Quite similar to Stereo A.
- SIR arrival: max difference = 2.6 days, mean = 14.0 hours, median = 11.0 hours
- Slow wind speed seems to be underestimated by J-map fits...
- ...but density peaks are not so relevant for slow wind (stream interface).
Need to measure 1-2 days before. (see comparison to L. Jian results).

- J-map running difference images track mostly the stream interface.
 - Several classes of events (CMEs, transitory overdensities...) can mask or mimic a SIR. Need the in-situ diagnostic.
 - When well identified SIR the arrival time is about 12 hours precise on Stereo A and B.
 - The SIR speed seems to be underestimated, but not so much if compare to the pre-stream interface in-situ SW speed.
- Complete non-seen SIR in J-maps.
 - 2009-2014 to be done. Going to the Solar maximum...
 - Specific study on transitory events impact ?

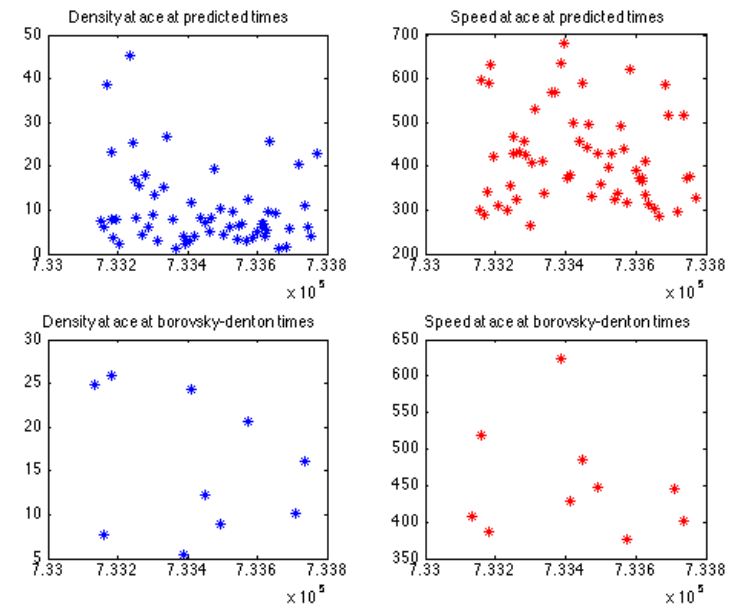
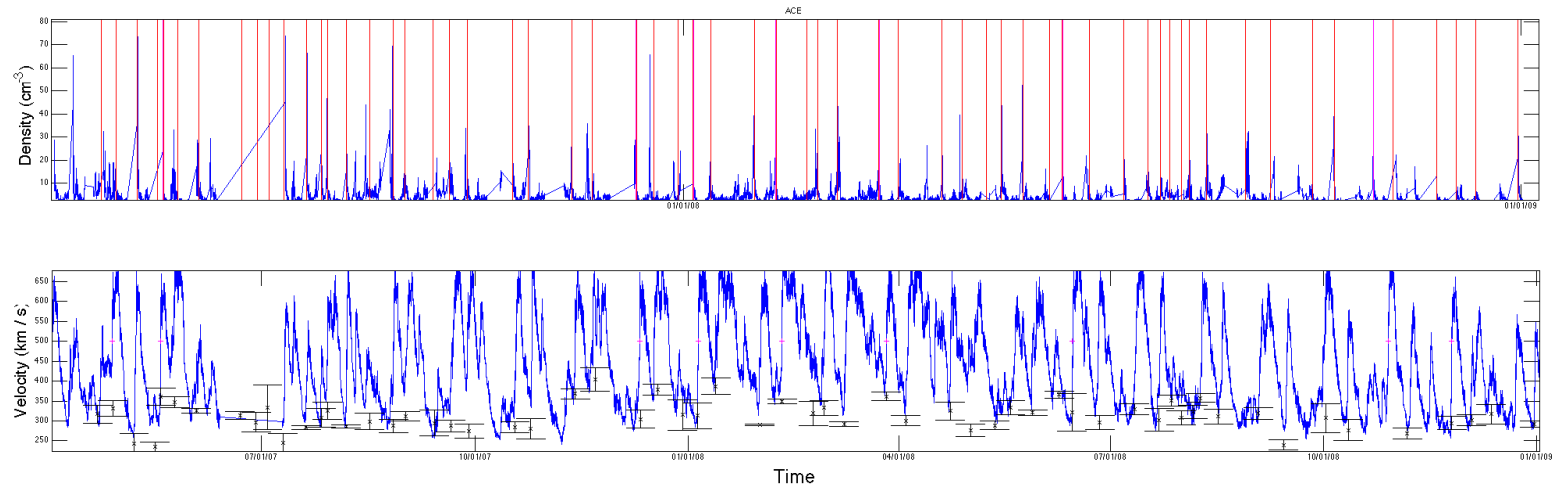
Extras

L.K. Jian, C.T. Russell, J.G. Luhmann, A.B. Galvin, K.D.C. Simunac,
Solar Wind Observations at STEREO: 2007 – 2011,
Amer. Inst. Phys. Proceedings of Solar Wind 13, 1539, 191-194, doi: 10.1063/1.4811020, 2013
http://www-ssc.igpp.ucla.edu/forms/stereo/stereo_level_3.html

- 248 events on StA and 231 events on StB in the period 2007-2014. ~ 30 /year
- Minimal speed: $\langle V_{\min} \rangle = 320$ km/s (312 from J-maps up to end of 2008)
- Maximal speed: $\langle V_{\max} \rangle = 570$ km/s



ACE measurements



- \rightarrow Compare to the Borovsky & Denton CIRs (2007-2008 available)